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# The relationship between total solar radiation and biologically erythematic radiation over urban region of Egypt



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#### ABSTRACT

This work studies the relationship between the effective of erythematic radiation (EER) and the total solar radiation (G) from urban region of Egypt. The measurements of total solar radiation (G) and biologically effective erythematic radiation (EER) incident on a horizontal surface at Cairo, Egypt (lat. 30°05′N and Long. 31°15′E), during the period time (1990–2010) are used. The relationship between hourly mean and daily mean of EER and the hourly mean daily mean totals of broadband total solar radiation is presented. The average hourly monthly mean variation of slant ozone Z and UVB transmission  $K_{\text{HJVB}}$  at the present work are found. The relation between the two variables slant ozone Z and UVB transmission  $K_{\text{tUVB}}$  show an opposite seasonal behavior, the low values of the slant ozone column during summer time produce high UVB transmission values in this season. The apposite pattern is observed during the winter. The Seasonal Statistical values of regression equations; the slopes  $(\beta)$ , intercepts  $(\alpha)$  and the standard errors (SE) for the fitted lines. The minimum slope occurs in winter, indicating that the percentage reduction in EER at higher SZA is larger than G. The variations of the slopes ( $\beta$ ) during the course of the year ranges from a minimum of 0.2214 in winter to a maximum of 0.2914 in summer, the intercepts also show their minimum and maximum values in cold humid and hot dry months. The temporal variability of the percentage ratio of the total hourly mean daily erythema to total hourly mean daily broadband solar global irradiation (EER/G) is presented. The estimated values of UVB solar radiation a good agreement with the measured values of the UVB solar radiation, the difference between the estimated and measured values of UVB solar radiation varies from 2.3% to 3.6%. © 2014 Elsevier Ltd. All rights reserved.

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#### 1. Introduction

Stratospheric ozone is known to be the most important atmospheric factor determining clear sky UV-B radiation reaching the Earth's surface. The potential increase of UV-B exposure is the cause of mounting concern about the ozone layer. There are, however, other effects that influence the UV radiant energy transfer: cloud cover, aerosols, tropospheric ozone, and other gaseous pollutants. The relationships between various phenomena taking place in the atmosphere are complex and not well known. Therefore, ground based UV measurements are necessary to explore atmospheric changes and resultant effects on the biosphere. The spectral range of solar radiation corresponding to

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wavelengths  $\lambda$  < 400 nm, is called ultraviolet (UV). The ultraviolet is subdivided into three wavelength band regions: the UVA (315-400 nm) which is received at earth's surface, the UVB (280-315 nm) which is partially absorbed by ozone or scattered in the atmosphere and UVC (<290 nm) which is potentially the most dangerous as it has the highest energy levels, but this wavelength band region is completely absorbed by stratospheric ozone and oxygen above about 30 km. The knowledge of solar UV radiation reaching the Earth's surface has a great interest because of its significant role in atmospheric and biological processes. The UV-B solar radiation (280-315 nm) represents less than 1% of the total radiation reaching the earth surface, and it is very important for the Earth-living systems because it is a radiation of high energy. UV-B irradiance on the Earth surface depends on geographical factors such as latitude, height, earth-sun distance, and solar zenith angle (SZA), etc. The influence of these factors can be evaluated using different radiative models. However UV-B solar radiation depends on atmospheric parameters like ozone, clouds and aerosols. Ozone is the gas that absorbs UV-C and some UV-B solar radiation and the effect of the total ozone column is included in all radiative models. Clouds are another attenuating factor of UV-B radiation and due to their random nature they are difficult to model. Aerosol is the factor that affects radiation levels under cloudless sky conditions [1-3].

During the last years, concerns about the intensity levels of UV-B radiation reaching the ground have increased due to the stratospheric ozone depletion and the dramatic increase in the number of skin cancers in the population. The UV solar irradiance at the ground varies greatly with local time, latitude and season, primarily because of the changing elevation of the sun in the sky. The ozone plays a role of shield around the Earth protecting us from ultraviolet radiation. The UVB radiation only represents 5% of the UV radiation [4], and 0.5% of the solar radiation, but UVB is very important to human beings because it can produce different illnesses [5]. Frequently the biologically effective irradiance is given as UV index (UVI). UVI is a dimensionless quantity and one unit is equivalent to 25 mW m<sup>-2</sup> of erythematic radiation. The Earth's atmosphere significantly modifies the incoming solar radiation through the absorption and scattering process by gases, dust particles and other biosphere constituents of human and natural activities, there is clear linear correlation between UV-B total and global solar radiation, particularly in the region of moderate to low global irradiance, which enables estimation of UV-B flux in tropical/equatorial areas where facilities for UV-B measurements are not available, but global solar radiation flux data exist [3,6-9].

The relationship between total ozone and spectral UV irradiance from Brewer spectrophotometer observations and its use for derivation of total ozone from UV measurements has been the subject of [10]. The most significant influence on received clear sky UVB is that resulting from variation in atmospheric ozone [11]. Much of the harmful UV-B is absorbed by stratospheric ozone, although downward trends observed in total column ozone, particularly at high latitudes, and to a lesser extent at midlatitudes, imply significant increase in the surface UV exposure [12]. The anticorrelation between total column ozone and UV radiation is a complex function of many variables, as: solar zenith

angle, solar elevation, cloud cover, aerosol and vertical profile of ozone. The attenuation of UVB radiation by clouds is frequently larger than any other atmospheric parameter; however it is often only approximated by modeling the effect of monthly or seasonally averaged cloud amounts [13]. Ground-based observations can play an important role in improving the understanding of some of these effects [14,15].

The beneficial effects of UV-A and UV-B radiation on humans, the ecosystem, animals, plants, and materials have been addressed by many investigators [3,16-20]. The study of the ratios of biological UV to G have received a considerable attention in the past few years so that relationships of this type have been proposed by different investigators with measurements from Kuwait [21], Dharan [22], Edmonton [23], Saudi Arabia [24], Valencia [25,26], Spain [27,28], Egypt [29], France [30], Iran [31] and Switzerland [32]. In another study a relationship was established between an 18-month record of daily UV-B and full band (300-3000 nm) solar radiation in Sutton (England) [33]. They suggested a linear relation between the ratio of energy in the two wavebands and the cosine of zenith angle at noon time which enables the UV-B irradiation to be estimated from full band (G) solar radiation. Using a series of measurements including ultraviolet UV-B (280-320 nm), UV-A (320-400 nm), and broadband global (250-2800 nm) made from June 1998 to August 2001 at a station in Kwangju, South Korea, Ogunjobi and Kim [34] concluded that the ratio of total UV (280-400 nm) to broadband radiation is about 7.7% for all-sky conditions. The ratio of the ultraviolet to global radiation (UV/G) was also calculated by [19], for two cities in Egypt and compared with other sites in the Arabian Peninsula. Part of this study was used before to evaluate the prediction of clear-sky biologically effective erythematic radiation from global solar radiation [3].

In the present work, the estimation of hourly mean daily erythema (EER) from hourly mean daily broadband (G) in urban regions has rarely been studied, the main aim of this study is to establish an empirical relationship between total solar radiation (G) and effective erythema radiation (EER), which enables the estimation of the daily integrated EER from easily available broadband (250–2800 nm) G data in such climates, and the data in this study were obtained from the Meteorological Authority of Egypt.

#### 2. Methodology and model variables

UVB measurements were converted into UVER (solar ultraviolet erythematic irradiance) values by means of conversion factors, (Diffey factor) provided by the manufacturer [35], and from them UVI (ultraviolet indices) hourly values were evaluated [36]. These UVI results have been considered as measured values and represented by (UVI) mean values (Middle East airlines values). UVI values have also been obtained from spectral calculated weighted by the erythema action spectrum, they are represented by (UVI) model and can be obtained by the following expression:

$$(UVI)_{\text{model}} = K_{\text{er}} \int_{290}^{400} E_{\lambda} S_{\text{er}} (\lambda) d\lambda$$
 (1)

**Table 1**Monthly mean values of extraterrestrial solar radiation (Gmax.), long-term 1990–2010, total solar radiation (G) and air mass in noon time (m) at Cairo, Egypt.

Month	Jan.	Feb.	Mar.	Apr.	May	Jun.	Jul.	Aug.	Sep.	Oct.	Nov.	Dec.
$G_{\text{max}}$ . (Mj $^{m-2}$ d <sup>-1</sup> )	20.4	24.5	32.1	36.8	39.5	42.2	39.6	36.9	35.2	28.5	24.6	21.2
G (Mj $^{m-2}$ d <sup>-1</sup> )	12.8	14.9	16.8	19.5	25.3	27.6	26.4	24.2	21.6	19.3	14.9	13.7
m	1.47	1.31	1.15	1.06	1.03	1.01	1.04	1.06	1.11	1.18	1.31	1.42

**Table 2**The average hourly monthly mean variation of slant ozone and UVB transmission at Cairo during the period from January 1990 to December 2010.

		1990	1991	1992	1993	1994	1995	1996	1997	1998	1999	2000	2001	2002	2003	2004	2005	2006	2007	2008	2009	2010
J	$\begin{matrix} K_t \\ Z \end{matrix}$	0.062 415	0.052 429	0.069 412	0.058 385	0.071 394	0.064 410	0.055 419	0.074 428	0.063 439	0.065 416	0.057 389	0.067 402	0.061 388	0.077 415	0.054 396	0.072 405	0.064 419	0.058 422	0.062 397	0.070 408	0.065 397
F	$\begin{matrix} K_t \\ Z \end{matrix}$	0.071 391	0.076 385	0.084 394	0.072 377	0.088 381	0.073 369	0.068 376	0.081 385	0.079 391	0.086 389	0.079 381	0.074 377	0.075 384	0.086 379	0.067 385	0.087 391	0.084 375	0.069 379	0.074 384	0.083 375	0.077 381
M	$\begin{matrix} K_t \\ Z \end{matrix}$	0.091 342	0.098 355	0.088 348	0.096 339	0.0104 357	0.0109 345	0.094 351	0.096 357	0.087 339	0.094 344	0.093 357	0.0104 349	0.0108 338	0.0101 358	0.097 350	0.091 341	0.095 345	0.096 338	0.094 347	0.097 342	0.092 348
Α	$\begin{matrix} K_t \\ Z \end{matrix}$	0.112 300	0.117 305	0.0113 298	0.115 307	0.119 311	0.113 302	0.119 295	0.108 306	0.109 309	0.115 300	0.117 289	0.114 297	0.108 304	0.111 307	0.116 310	0.113 305	0.112 296	0.115 312	0.109 304	0.114 298	0.117 301
M	$\begin{matrix} K_t \\ Z \end{matrix}$	0.124 285	0.129 291	0.135 286	0.127 291	0.134 288	0.119 284	0.125 291	0.128 295	0.134 290	0.131 285	0.127 289	0.122 294	0.117 291	0.124 296	0.129 292	0.124 289	0.118 295	0.126 291	0.122 294	0.118 291	0.127 288
J	K <sub>t</sub> Z	0.131 270	0.136 275	0.139 281	0.133 277	0.137 275	0.141 280	0.138 283	0.142 276	0.135 281	0.138 277	0.144 271	0.139 274	0.141 274	0.137 277	0.139 279	0.142 281	0.138 274	0.140 276	0.137 280	0.141 284	0.135 276
J	$\begin{matrix} K_t \\ Z \end{matrix}$	0.127 282	0.130 285	0.127 288	0.129 291	0.127 293	0.130 289	0.125 287	0.127 286	0.131 290	0.124 282	0.122 284	0.128 287	0.125 293	0.122 291	0.129 286	0.124 282	0.126 284	0.127 289	0.128 291	0.122 286	0.124 284
Α	$\begin{matrix} K_t \\ Z \end{matrix}$	0.115 287	0.118 292	0.112 290	0.117 289	0.114 295	0.118 293	0.119 292	0.116 290	0.114 294	0.117 286	0.115 289	0.117 290	0.116 294	0.118 295	0.114 291	0.112 293	0.114 288	0.118 287	0.115 290	0.116 294	0.113 290
S	K <sub>t</sub> Z	0.102 322	0.105 325	0.103 329	0.099 319	0.102 322	0.097 327	0.096 326	0.101 318	0.103 315	0.105 325	0.102 328	0.096 322	0.099 317	0.102 324	0.106 328	0.101 324	0.098 322	0.096 319	0.103 320	0.104 327	0.097 329
0	$\begin{matrix} K_t \\ Z \end{matrix}$	0.091 351	0.095 355	0.088 359	0.091 348	0.087 345	0.091 344	0.086 350	0.084 356	0.090 348	0.094 350	0.096 354	0.089 348	0.094 352	0.091 357	0.093 351	0.087 347	0.089 356	0.094 357	0.095 352	0.094 348	0.092 354
N	K <sub>t</sub> Z	0.084 365	0.086 371	0.078 375	0.075 379	0.080 374	0.082 372	0.074 368	0.077 370	0.084 374	0.079 376	0.077 372	0.082 369	0.085 374	0.079 370	0.080 372	0.083 375	0.077 370	0.074 373	0.079 369	0.082 374	0.085 371
D	$\begin{matrix} K_t \\ Z \end{matrix}$	0.077 392	0.072 396	0.069 391	0.071 389	0.067 394	0.070 397	0.067 401	0.072 404	0.074 398	0.071 394	0.069 397	0.072 395	0.068 391	0.070 388	0.073 394	0.075 398	0.068 401	0.067 398	0.069 396	0.071 392	0.074 401

where  $K_{\rm er}$  is 40 m<sup>2</sup> W<sup>-1</sup>,  $E_{\lambda}$  is the UV spectrum wavelength dependent (Wm<sup>-2</sup> nm<sup>-1</sup>) and  $K_{\rm er}$  is the erythemal weighting function accepted by CIE (Commission International d'Eclairage) and given by [37].

The clearness index is chosen for characterizing jointly the cloud cover and aerosol load and it is evaluated by the following expression:

$$C.I. = \left[ \left\{ \frac{(D_h + I_n)}{D_h} + KZ^3 \right\} / \left\{ 1 + KZ^3 \right\} \right]$$
 (2)

where  $D_h$  is the horizontal diffuse irradiance,  $I_n$  is the normal incidence direct irradiance, Z is the solar zenith angle and k is a constant whose value is 1.041 with Z in radians, by means expression (2), the dependence between the variable, C.I., and the solar zenith angle has been removed [38].

The slant total ozone column, Dobson (DU) represents the actual ozone amount in the atmosphere that the solar radiation passes through [39–41], have defined it as follows:

$$Z = TOC/\mu \tag{3}$$

where  $\mu$  is the cosine of the solar zenith angle, this expression is only valid for the direct solar irradiance. However, it can be used as a good approximation for the global solar irradiance (direct+diffuse) since the largest part of ozone absorption occurs at high altitudes, before the scattering process by aerosol and cloud.

In analog with the broadband and UV cases, the UVB hemispherical transmittance can be defining in the following way [42,43].

$$K_{\text{TUV}} = \frac{\text{UVB}}{\text{UVBext.}} \tag{4}$$

where UVBext. is the extraterrestrial UVB radiation value on a horizontal surface it is given by

UVBext. = 
$$I_{SCUVB}(12/\pi) E_0 \int_2^1 \sin(\theta) d$$
 (5)

where  $(\theta)$  is the solar elevation angle,  $E_0$  is the correction factor for the eccentricity of the Earth's orbit,  $\omega_i$  (i=1 and 2) is the solar our angle at the beginning of period and at the end of period, respectively, and  $I_{SCUVB}$  is the UVB solar constant (21.51 W m<sup>-2</sup>). It has been obtained from the spectral values given by [6,44].

The relationship between effective erythematic radiation (EER) and global solar radiation (G) as following Eq. (6) and several investigations [9,26,45,46], the linear relationship:

$$EER = \beta G + \alpha \tag{6}$$

where EER is the daily erythem irradiance, G is the total daily broadband global radiation,  $\beta$  is the slope of the linear relation, and  $\alpha$  is the intercept. Under clear-sky conditions, SZA and TCO appear to be the major determinant affecting the ratio of EER and G. Inverse relation is known between the ozone density in the atmosphere and the amount of UV reaching the earth's surface [47].

Other variables affecting by day-to-day fluctuation (aerosol, water vapor, air pressure) have smaller influence than SZA and TCO. In order to remove the ozone dependency from the relationship, the values of  $\beta$  were normalized to the 21-years (1990–2010) monthly mean climatological values of TCO using the following correction factor [48]:

$$C_{O_3} = \exp[-k_{O_3} (\Delta O_3) m_i]$$
 (7)

where ( $\Delta O_3$ ) is the monthly mean difference between TCO in given month (i) and long-term climatological value, ( $m_i$ ) is monthly mean air mass at noon time shows in Table 1, and ( $k_{O_3}$ )=3.546 (atm cm)<sup>-1</sup> is the mean ozone absorption coefficient for the wavelengths 300–316 nm.

The Run-test were done on the daily observed data (EER and G) to make sure that the data are homogeneous and the variations of daily observed EER and G are caused only by climatic influences and not by other sources of errors (e.g. systematic errors caused by

**Table 3** Means monthly and seasonal averages of the extraterrestrial UVB solar radiation, mesured and estimated UVB solar radiation and clearness index  $K_{\text{tUVB}}$  of UVB radiation at Cairo during the period 1990–2010.

Month	$\begin{array}{c} \text{UVB}_{\text{ext.}} \\ \text{(MJ m}^{-2} \ h^{-1}) \end{array}$	$\begin{array}{c} \text{UVB} \\ \text{(MJ } m^{-2} \ h^{-1} \text{)} \end{array}$	$\begin{array}{c} \text{UVB}_{est.} \\ \text{(MJ m}^{-2}  h^{-1}) \end{array}$	$K_{\text{tUVB}}$
Jan.	0.0755	0.0059	0.0042	0.071
Feb.	0.0774	0.0061	0.0055	0.082
Mar.	0.0816	0.0088	0.0072	0.101
Apr.	0.0828	0.0092	0.0086	0.118
May	0.0887	0.0107	0.0113	0.122
Jun.	0.0944	0.0125	0.0123	0.135
Jul.	0.0931	0.0133	0.0137	0.142
Aug.	0.0874	0.0121	0.0124	0.133
Sep.	0.0856	0.0100	0.0115	0.120
Oct.	0.0834	0.0087	0.0092	0.102
Nov.	0.0812	0.0068	0.0071	0.081
Dec.	0.0766	0.0052	0.0052	0.070
Winter	0.0782	0.0061	0.0065	0.084
Spring	0.0876	0.0114	0.0117	0.127
Summer	0.0883	0.0129	0.0128	0.134
Autumn	0.0809	0.0068	0.0071	0.084

**Table 4**Climatological monthly means of meteorological parameters at Cairo during the period 1990 to 2010.

Month	Total column ozone (Dobson) (TCO)	Ratio of bright sunshine hours (n/N)	Total number of clear sky days (NCD)	Total number of dust days (NDD)	Relative humidity (RH %)
Jan.	291.5	0.638	17.2	1.4	52
Feb.	297.2	0.681	15.2	1.6	50
Mar.	288.4	0.647	13.5	2.7	48
Apr.	319.6	0.632	12.1	3.2	44
May	313.2	0.722	15.8	3.7	36
Jun.	306.4	0.857	25.4	3.5	32
Jul.	292.6	0.871	28.6	2.6	28
Aug.	295.3	0.834	29.7	2.4	24
Sep.	284.7	0.841	27.5	2.7	30
Oct.	282.2	0.759	25.4	2.9	39
Nov.	271.3	0.727	20.3	2.2	47
Dec.	295.4	0.659	18.3	1.4	61

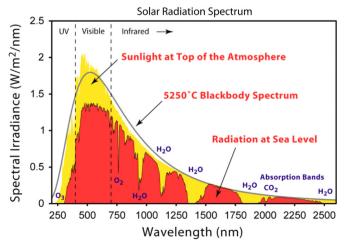


Fig. 1.

instruments; calibration problems; data transferring; etc.). These data are used to determine the temporal variability of the percentage ratio of the total daily erythema to total daily broadband solar global irradiation ( $\operatorname{EER}/G$ ). Note that, for calculating the annual cycle of the air mass (m) in Cairo, Egypt. The monthly mean solar zenith angle at noon time was used (Table 1). The analysis has been applied for clear-sky days when cloud amounts were less than two Oktas.

#### 3. Results and discussion

Table 2, shows the average hourly monthly mean variation of slant ozone Z and UVB transmission  $K_{\rm tUVB}$  at Cairo, Egypt during the period from January 1990 to December 2010. From this table, it is clear that the two variables show an opposite seasonal behavior, the low values of the slant ozone column during summer time produce high UVB transmission values in this season. The apposite pattern is observed during the winter. The average hourly monthly mean of slant ozone column and UVB transmission values shows the relationship between them in a clearer way than those of daily values. In winter months, UVB transmission values are low than those in summer months, this due to the fact that, the slant ozone column crossed by UVB radiation is higher in winter than in summer.

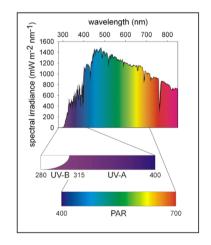
The mean monthly and the seasonal averages of the extraterrestrial UVB solar radiation measured and estimated UVB solar radiation and clearness index  $K_{\text{tUVB}}$  of UVB radiation at Cairo during the period time 1990–2010 are listed in Table 3. From this table we notice that, the maximum values of the above parameters occur around the summer months, while the minimum values in winter months. But the values of these variables is clear that, in the spring and autumn months fall between the values of the summer and winter months. And also clear that from Table 3, the estimated values of UVB solar radiation a good agreement with the measured values of the UVB solar radiation. The difference

between the estimated and measured values of UVB solar radiation varies from 2.3% to 3.6%. The average monthly UVB clearness index ( $K_{\rm tUVB}$ ) is less than corresponding values for global radiation  $k_{\rm t}$ . Where the value of  $k_{\rm t}$  is equal the value of global solar radiation (G) dividing by the value of extraterrestrial global solar radiation  $G_{\rm ext}$ . However, the  $K_{\rm tUVB}$  values vary from 0.070 to 0.142. This behavior is due to the extremely high attenuation of UVB radiation by stratospheric ozone and scattering phenomena.

The meteorological parameters (bright sunshine hours; relative humidity: total number of clear sky days: and total number of dusty days) have also been recorded during the present work (Table 4). To calculate the monthly mean values from the daily data, the work of Sabziparvar and Shetaee [48] was adapted. A dimensionless empirical relationship in the form of EER/G was developed, which can allow for the estimation of EER radiation from commonly measured global solar radiation (G). The daily clear-sky EER radiation can be estimated from broadband G data. Fig. 1-3 show the daily variability of clear-sky global solar radiation (G) and daily erythema radiation (EER), for the period of measurements. As seen, the daily EER is highly correlated with G, suggesting that daily EER doses can be obtained directly from G data. A linear regression line was fitted to the measured daily integrated G (as the independent variable) and daily EER (as the dependent variable). The regression analysis was applied to each season.

Table 5 performs the linear relations, the slopes  $(\beta)$ , intercepts  $(\alpha)$  and the standard errors (SE) for the fitted lines. As shown, the minimum slope occurs in winter, indicating that the percentage reduction in EER at higher SZA is larger than G. This can be explained by the increased ozone path length and higher percentage of scattering during the winter months.

In theory, the EER radiation must be zero when G is zero. As a result, the intercept  $(\alpha)$  of the relations might be expected to be zero. In this study, it was not the case in Table 5. This can be explained by the uncertain relative behavior of EER and G under darkness conditions. Furthermore, when solar zenith angles are



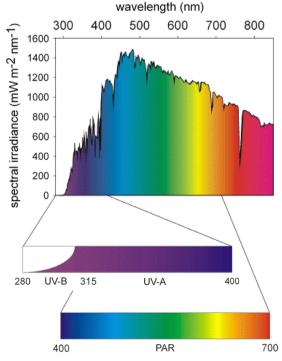


Fig. 2.



Fig. 3.

high (i.e. sunrise, sunset, winter months), the relative reduction in EER is disproportionately more than reduction in *G* because of the larger optical depths of ozone absorption and Rayleigh scattering.

As presented in Table 5, the variations of the slopes ( $\beta$ ) during the course of the year ranges from a minimum of 0.2214 in winter to a maximum of 0.2914 in summer. The intercepts also show their

minimum and maximum values in cold humid and hot dry months respectively. This result emphasis that the percentage reduction in the EER dose is always higher than such reduction in *G* when SZAs are high (i.e. winter). In addition to the seasonal relations, the following general relation was found to be reliable for estimation of daily integrated EER from global broadband solar radiation

**Table 5**The seasonal statistical values of regression equations during the period of the present work (1990–2010).

Season	Effective erythema radiation of equation (Kj m $^{-2}$ d $^{-1}$ ) (EER= $\beta G + \alpha$ )	Standard error of the slop of equation SE ( $\beta$ ) (Kj m <sup>-2</sup> d <sup>-1</sup> )	Standard error of the interception SE ( $\alpha$ ) (Kj m <sup>-2</sup> d <sup>-1</sup> )	Coefficient of determination $(R^2)$	Correction factor for the removal of ozone effect $(C_{0_3})$
Winter	$EER = 2.214 \times 10^{-1} \text{ G} - 0.685$	± 0.0137	± 0.087	0.939	0.93
Spring	$EER = 2.845 \times 10^{-1} \text{ G} - 1.258$	$\pm 0.0154$	$\pm  0.154$	0.885	0.96
Summer	$EER = 2.914 \times 10^{-1} \text{ G} - 2.332$	$\pm 0.0124$	$\pm  0.235$	0.915	0.94
Autumn	$EER = 2.521 \times 10^{-1} \text{ G} - 2.185$	$\pm0.0081$	$\pm 0.059$	0.954	0.91

throughout the year:

EER (Kj m<sup>-2</sup> d<sup>-1</sup>) = 
$$2.268 \times 10^{-1}$$
G (Mj m<sup>-2</sup> d<sup>-1</sup>)  $- 2.411$  (8)

It should be reminded that other variables (aerosol, water vapor, and haze) affecting the two wave bands by day-to-day fluctuation, have a smaller influence than the annual SZA and TCO cycles. Additionally, the disproportional effects of the surface albedo and dust aerosols on EER and *G*, might explain some of the contradictions observed in the linear relations. Improved correlations and less dependence on location can be achieved by using longer period data at other latitudes and all weather conditions.

#### 4. Conclusion

In this research we studies the average hourly monthly mean variation of slant ozone Z and UVB transmission  $K_{\rm tUVB}$  at Cairo, Egypt during the period from January 1990 to December 2010. The two variables show an opposite seasonal behavior, the minimum values of the slant ozone column during summer time produce high UVB transmission values in this season. The apposite pattern is observed during the winter. The average hourly monthly mean of slant ozone column and UVB transmission values shows the relationship between them in a clearer way than those of daily values. In winter months, UVB transmission values are low than those in summer months, this due to the fact that, the slant ozone column crossed by UVB radiation is higher in winter than in summer.

The prediction values of UVB solar radiation a good agreement with the measured values of the UVB solar radiation, the difference between the prediction and measured values of UVB solar radiation varies from 2.3% to 3.6%. The Seasonal Statistical values of regression equations; the slopes  $(\beta)$ , intercepts  $(\alpha)$  and the standard errors (SE) for the fitted lines. As shown, the minimum slope occurs in winter, indicating that the percentage reduction in EER at higher SZA is larger than G. This can be explained by the increased ozone path length and higher percentage of scattering during the winter months. The variations of the slopes  $(\beta)$  during the course of the year ranges from a minimum of 0.2214 in winter to a maximum of 0.2914 in summer, the intercepts also show their minimum and maximum values in cold humid and hot dry months respectively. In general, the good relation was found to be reliable for estimation of daily integrated EER from global broadband solar radiation throughout the year as following:

EER (Kj m<sup>-2</sup> d<sup>-1</sup>) = 
$$2.268 \times 10^{-1}$$
 G (Mj m<sup>-2</sup> d<sup>-1</sup>) -  $2.411$ 

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